

Interstellar astrometric navigation

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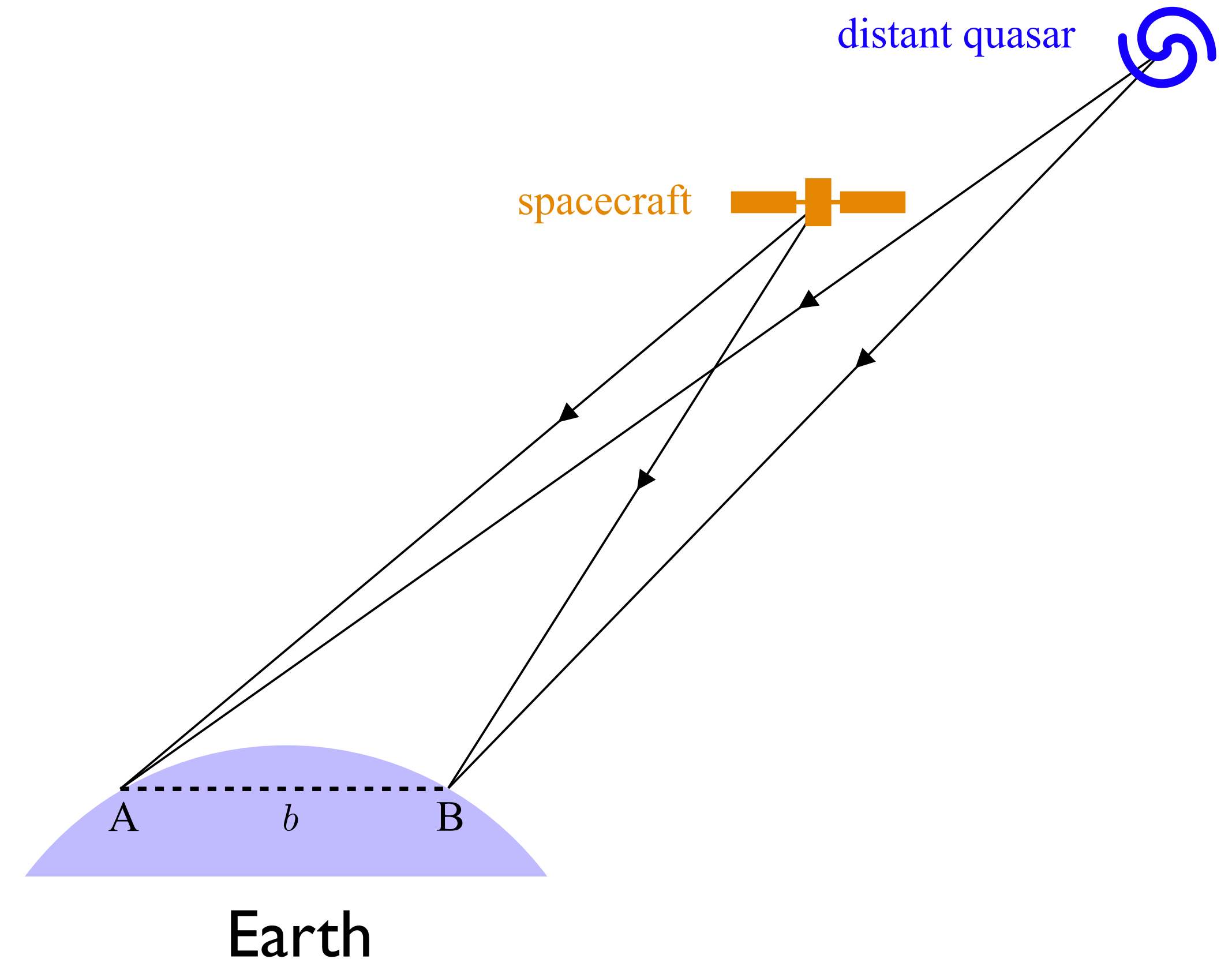
*Image source: Samuel Walters (1857)
from wikimedia commons (public domain)*



Navigation requirements

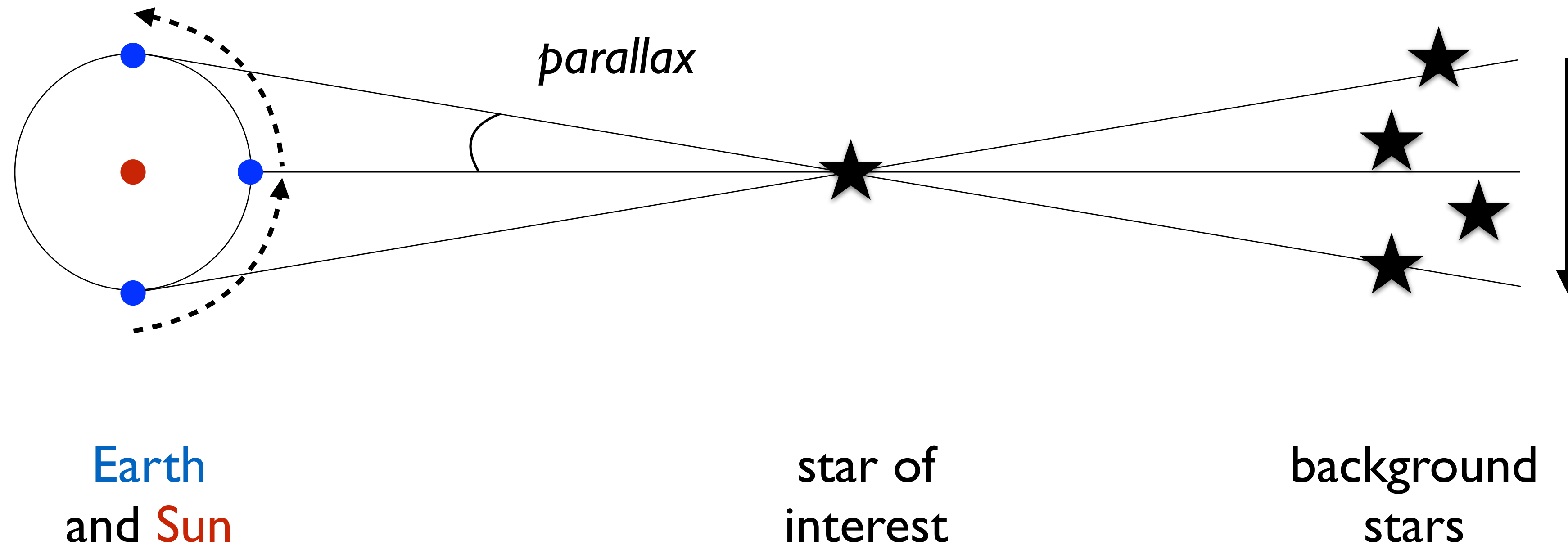


- On-board position and velocity required for
 - determining and correcting trajectory
 - pointing instruments and the antenna
- Navigation in the solar system
 - inertial: accelerometers, gyroscopes, clocks
 - Earth-based ranging, e.g. delta-DOR
- Interstellar navigation
 - must be entirely autonomous



Interplanetary navigation using delta-DOR

Parallax



- By definition, a star 1 pc (3.26 ly) away has a parallax of 1 arcsecond ($5 \mu\text{rad}$)
- Turn the use of parallax around: knowing distance, determine change in position

Aberration



- Also for photons: consequence of relative motion and the finite speed of light
- Position of stars change our velocity relative to them changes

Aberration

- Apparent position of any star from Earth varies over the year by $\pm 21''$ (30 km/s)

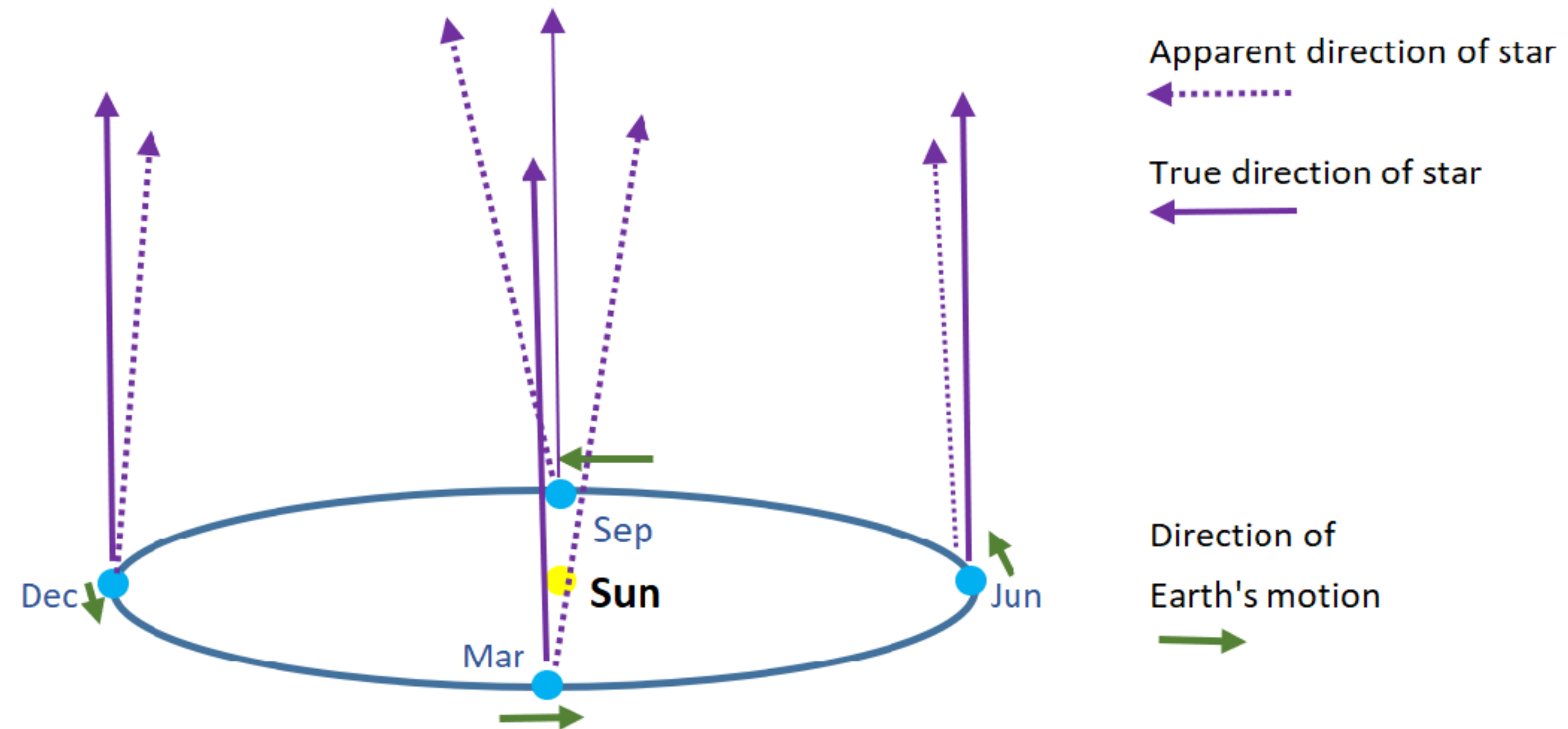
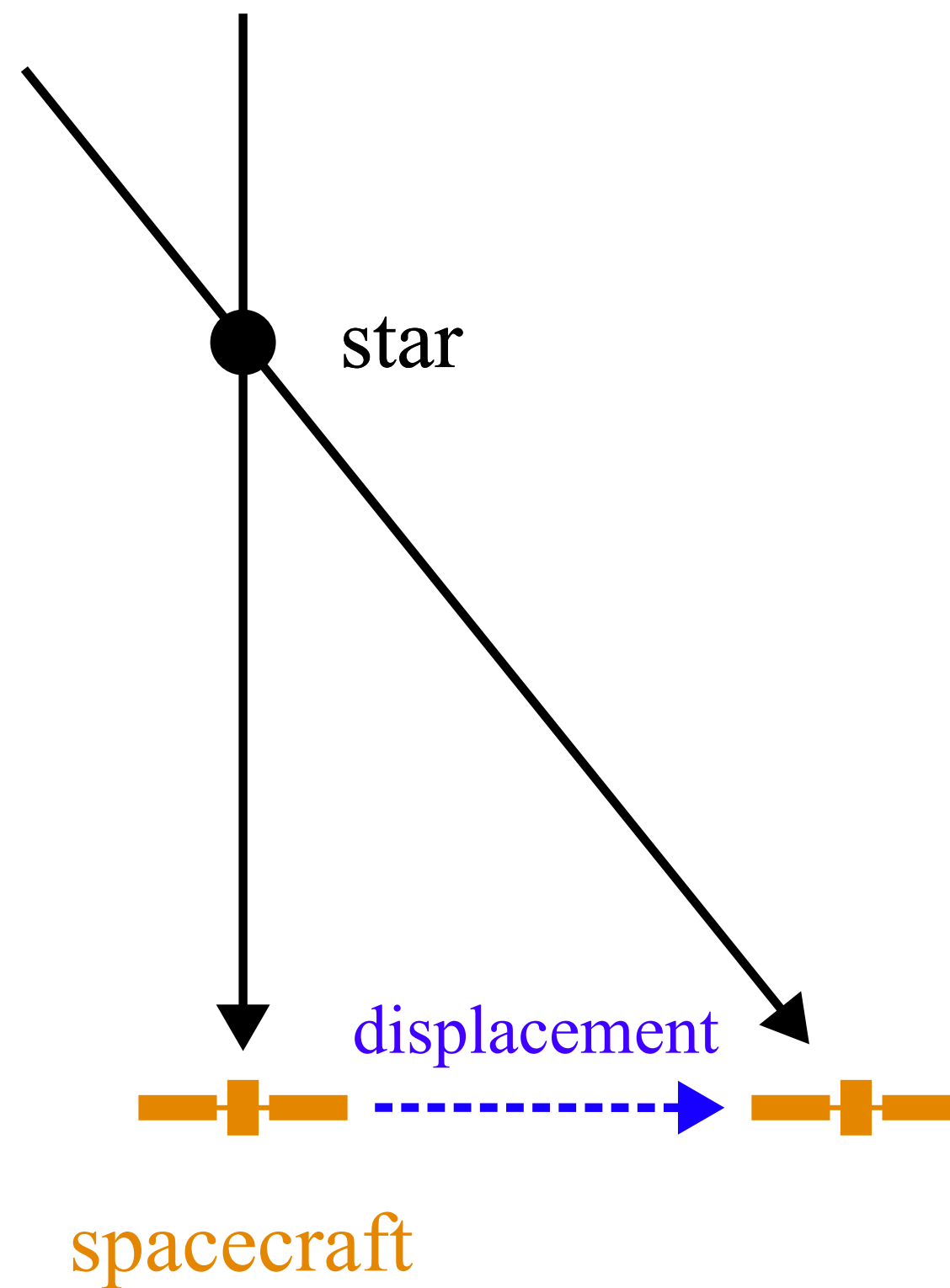


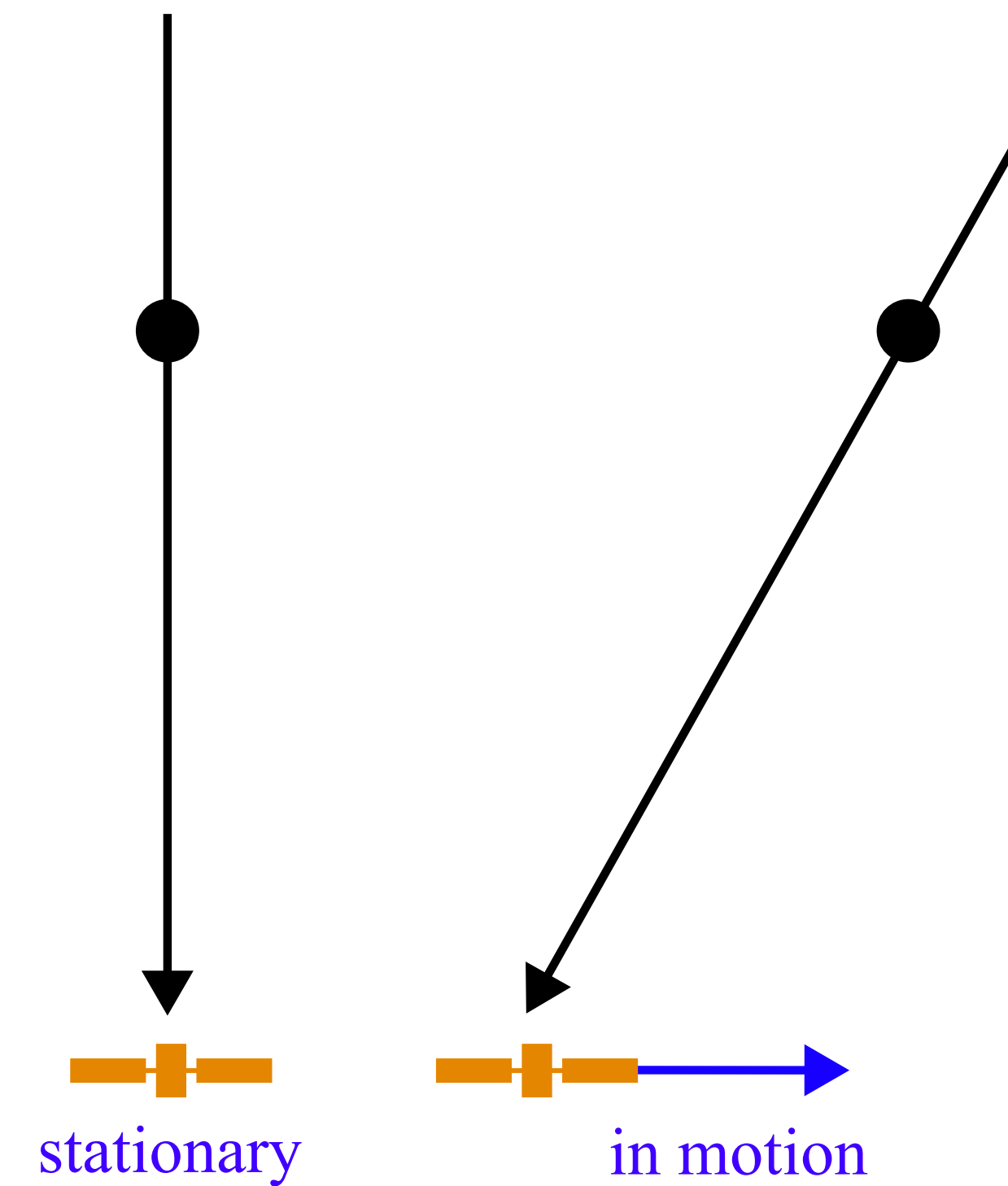
Image source: <https://thesciencegeek01.files.wordpress.com>

Combined parallax and aberration

shift due to parallax



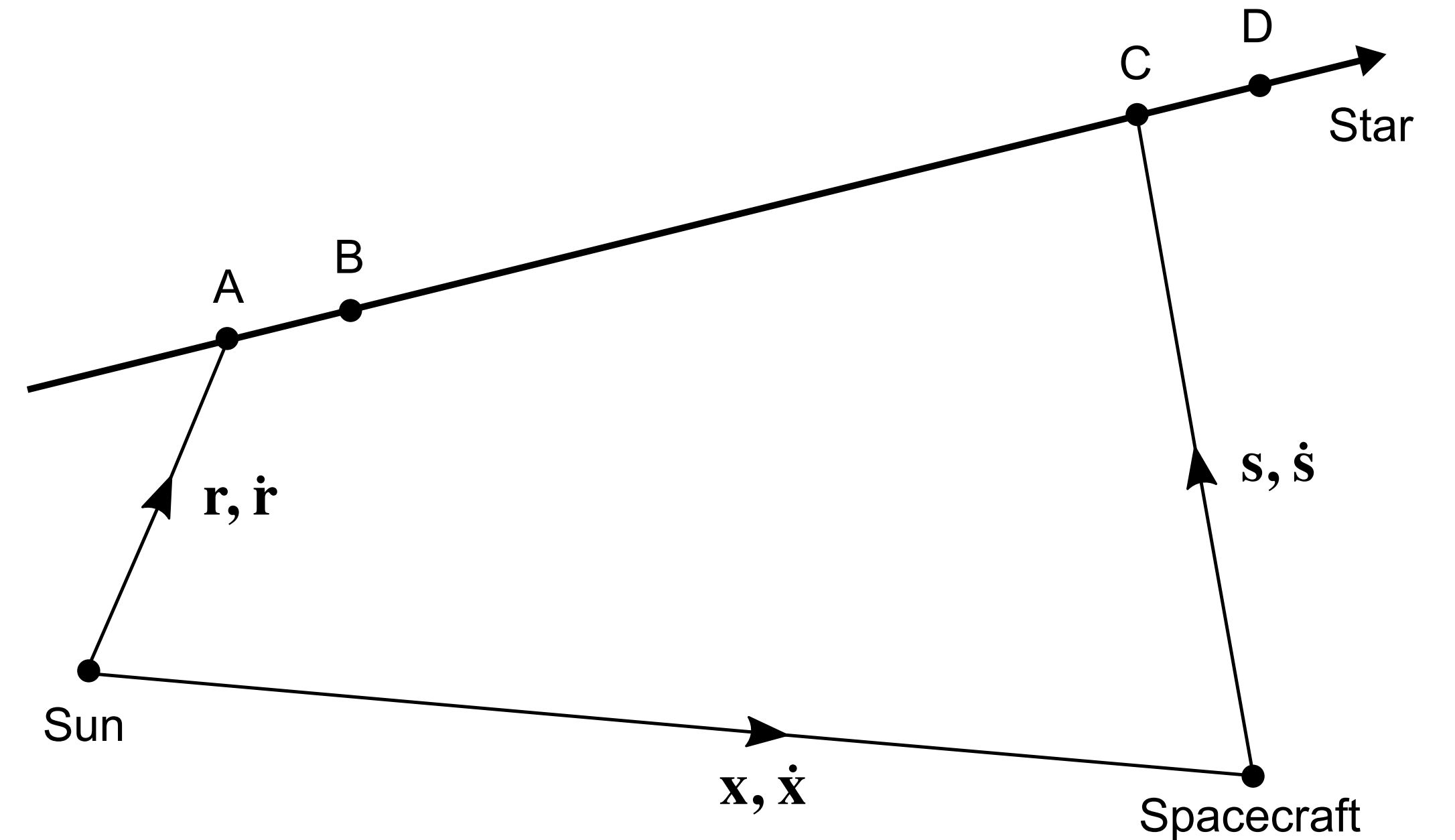
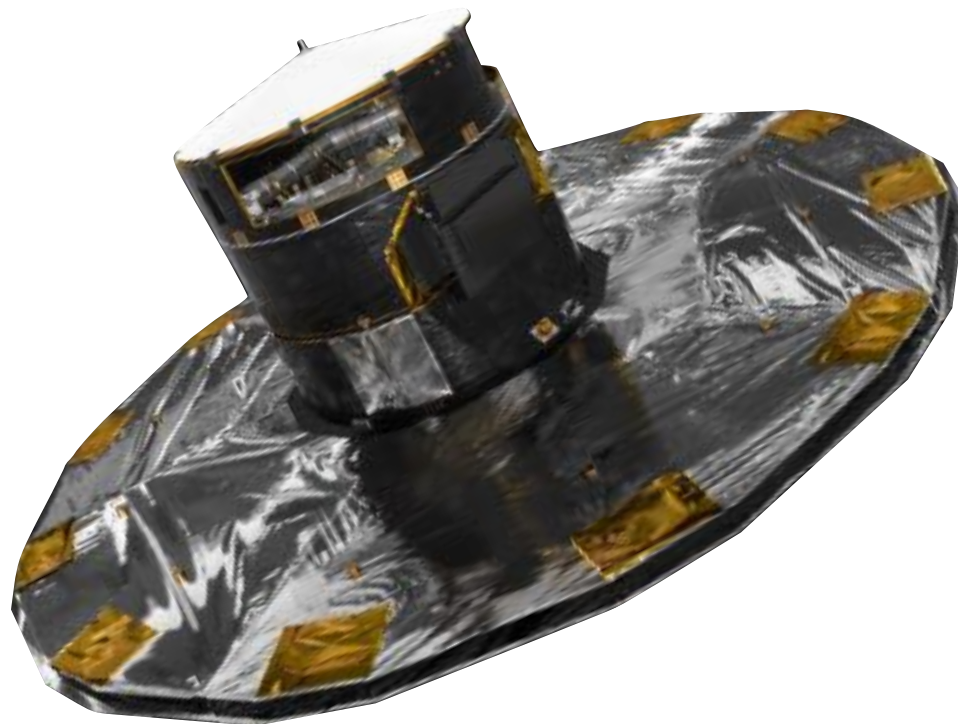
shift due to aberration



Stellar astrometric navigation

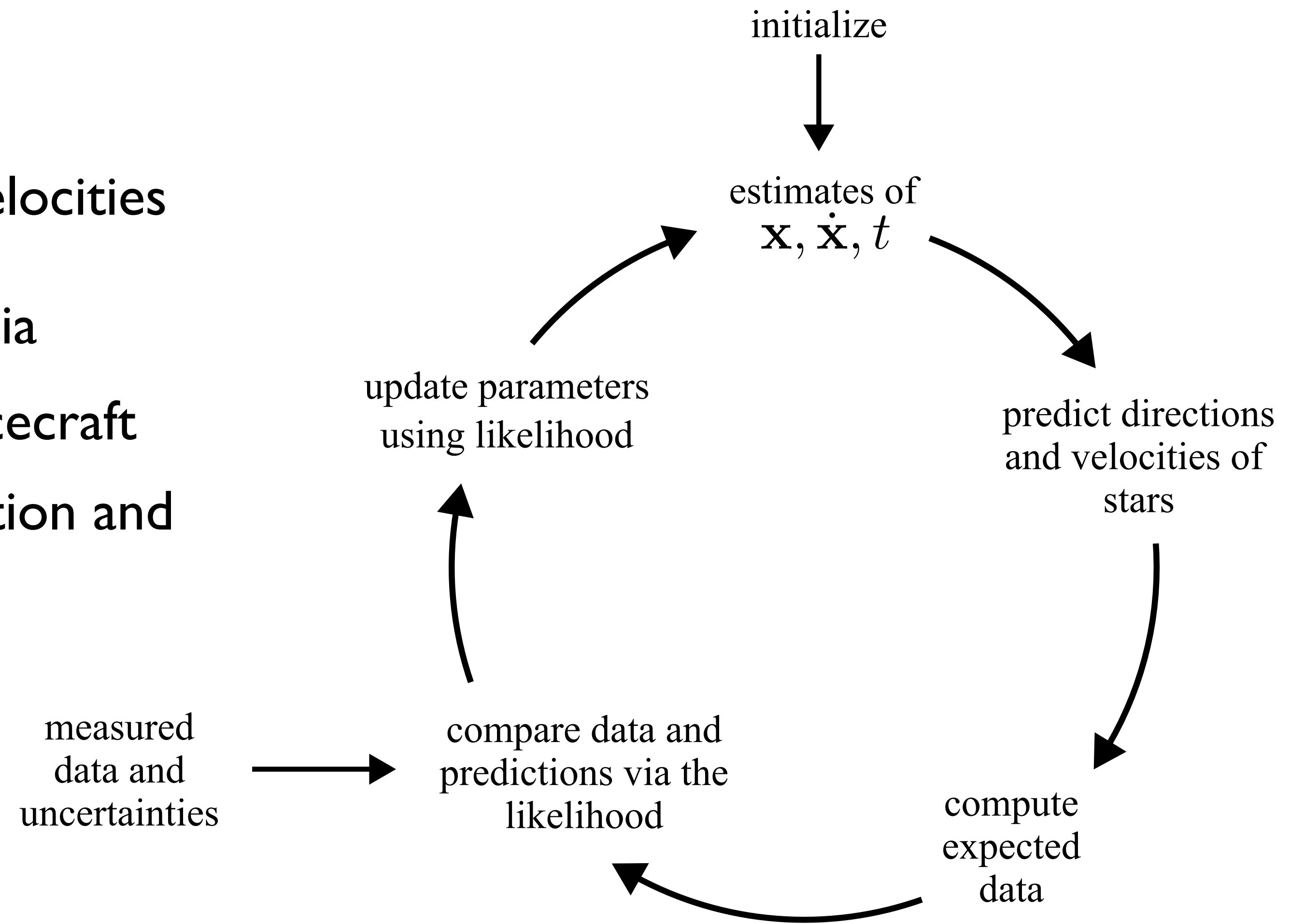


- Stars also move relative to the Sun
- To navigate, need to know 3D positions and 3D velocities of stars
- Need multiple stars to untangle effects of parallax and aberration
- ESA's Gaia Galactic survey mission



Iterative solution

- We have
 - catalogue of 3D positions & 3D velocities of stars relative to Solar System Barycentre (SSB, $\approx$ Sun) - from Gaia
 - measured star positions from spacecraft
 - initial estimates of spacecraft position and velocity ($\mathbf{x}, \dot{\mathbf{x}}$)
- Bayesian Monte Carlo



Spacecraft measurements

- Cannot measure angular coordinates (e.g. Right Ascension and Declination) accurately in a local coordinate system
- So measure angular separations (θ) between pairs of stars

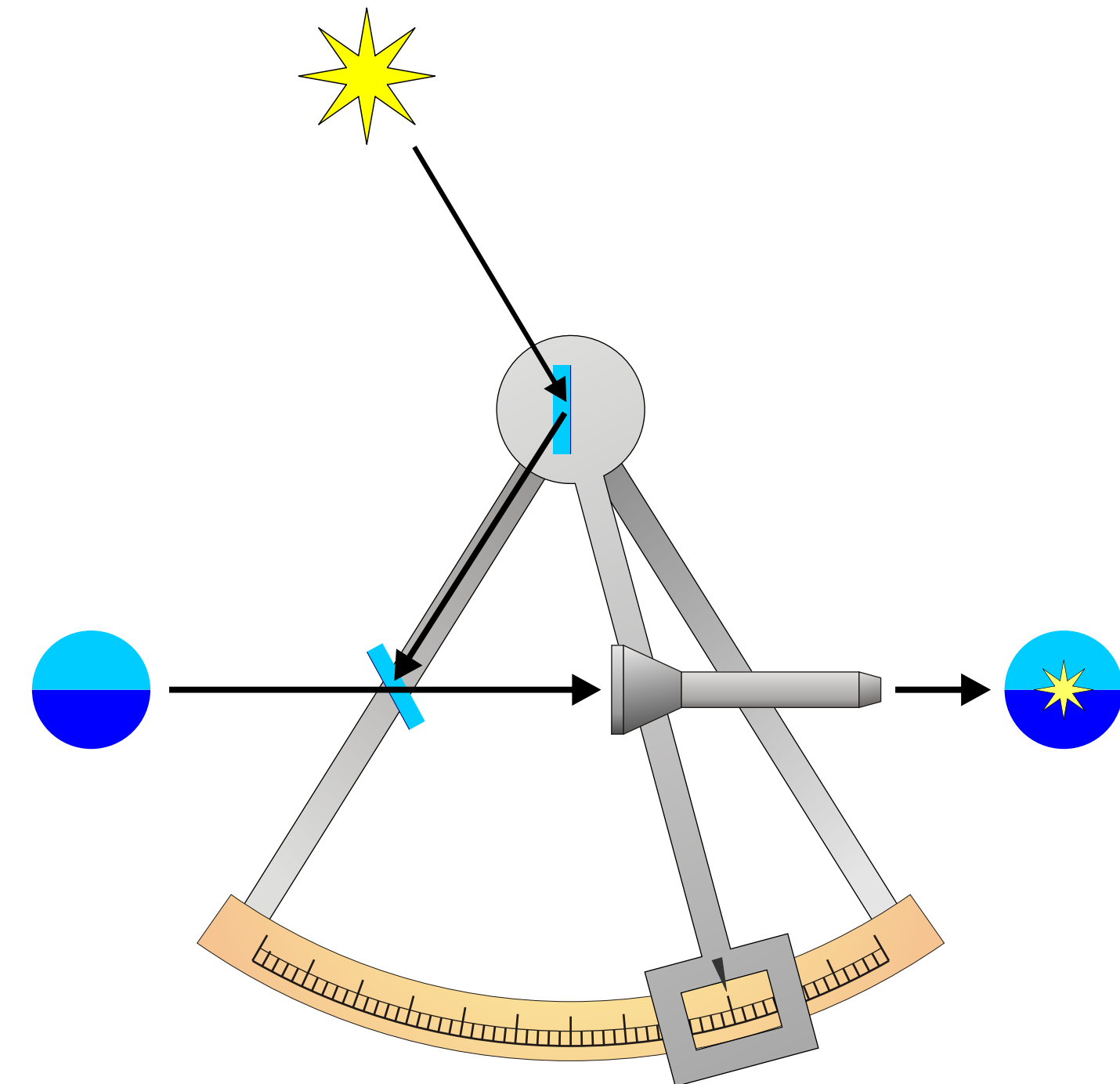
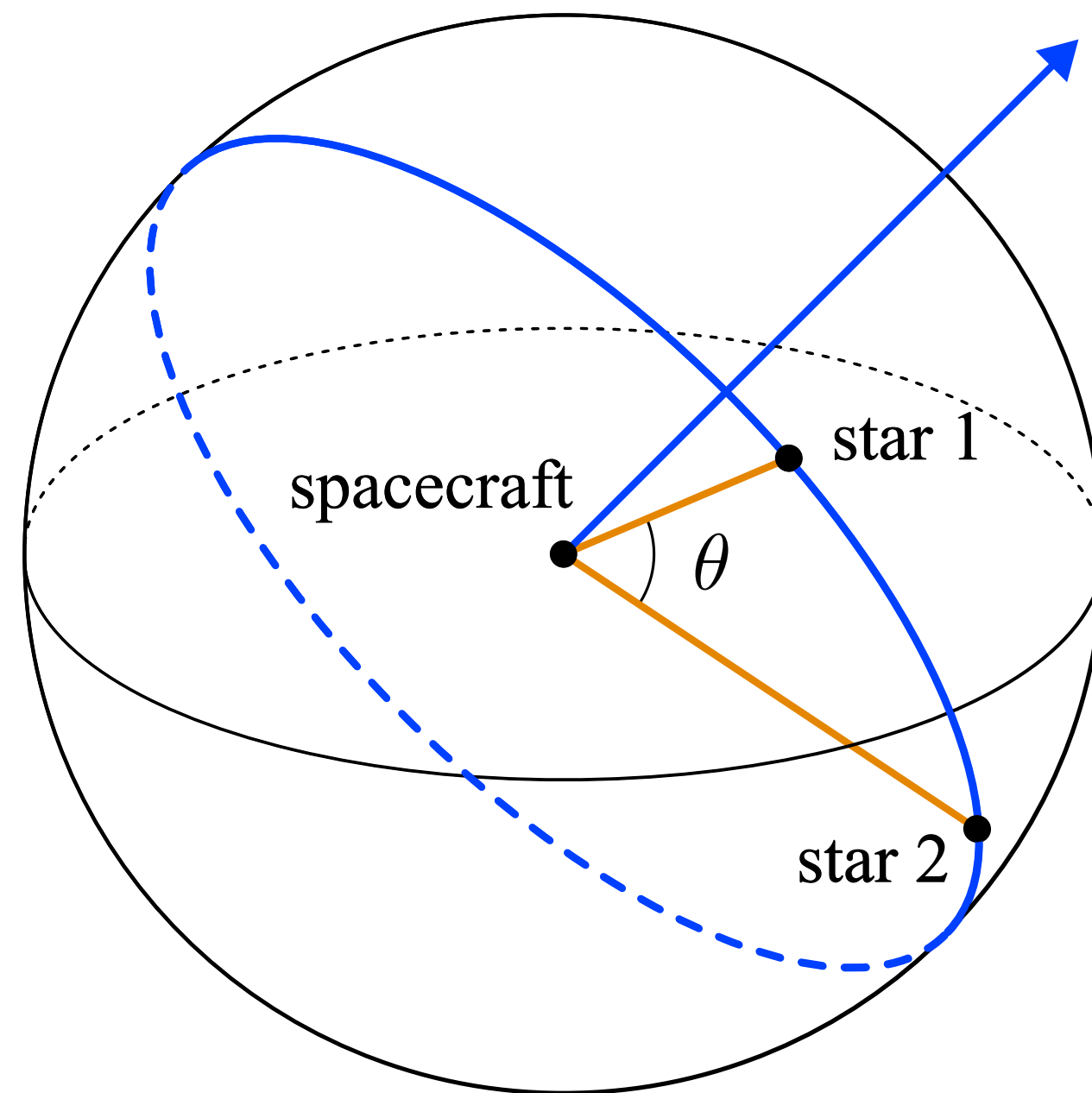


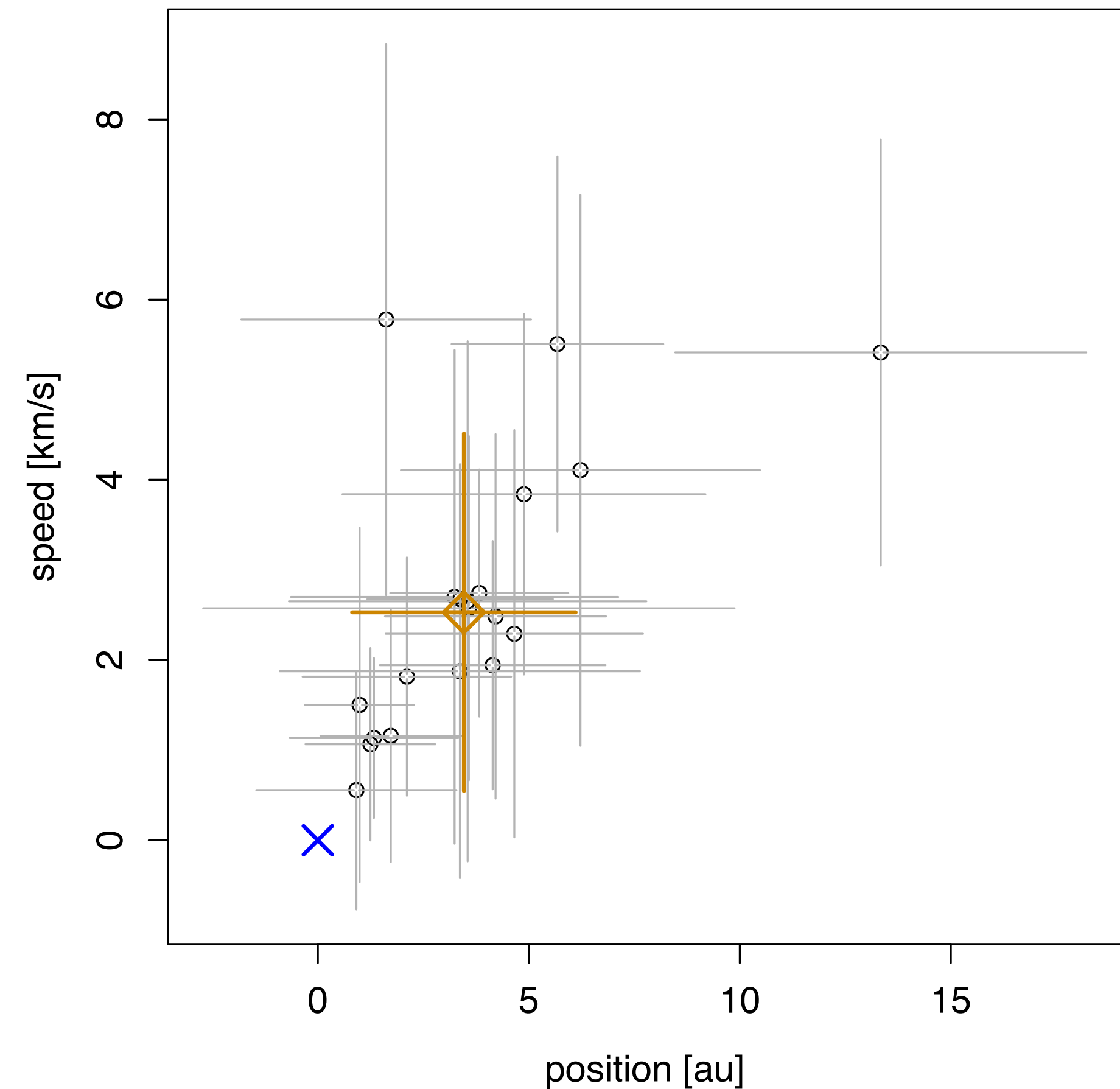
Image source (right): modified from wikipedia commons (CC0)

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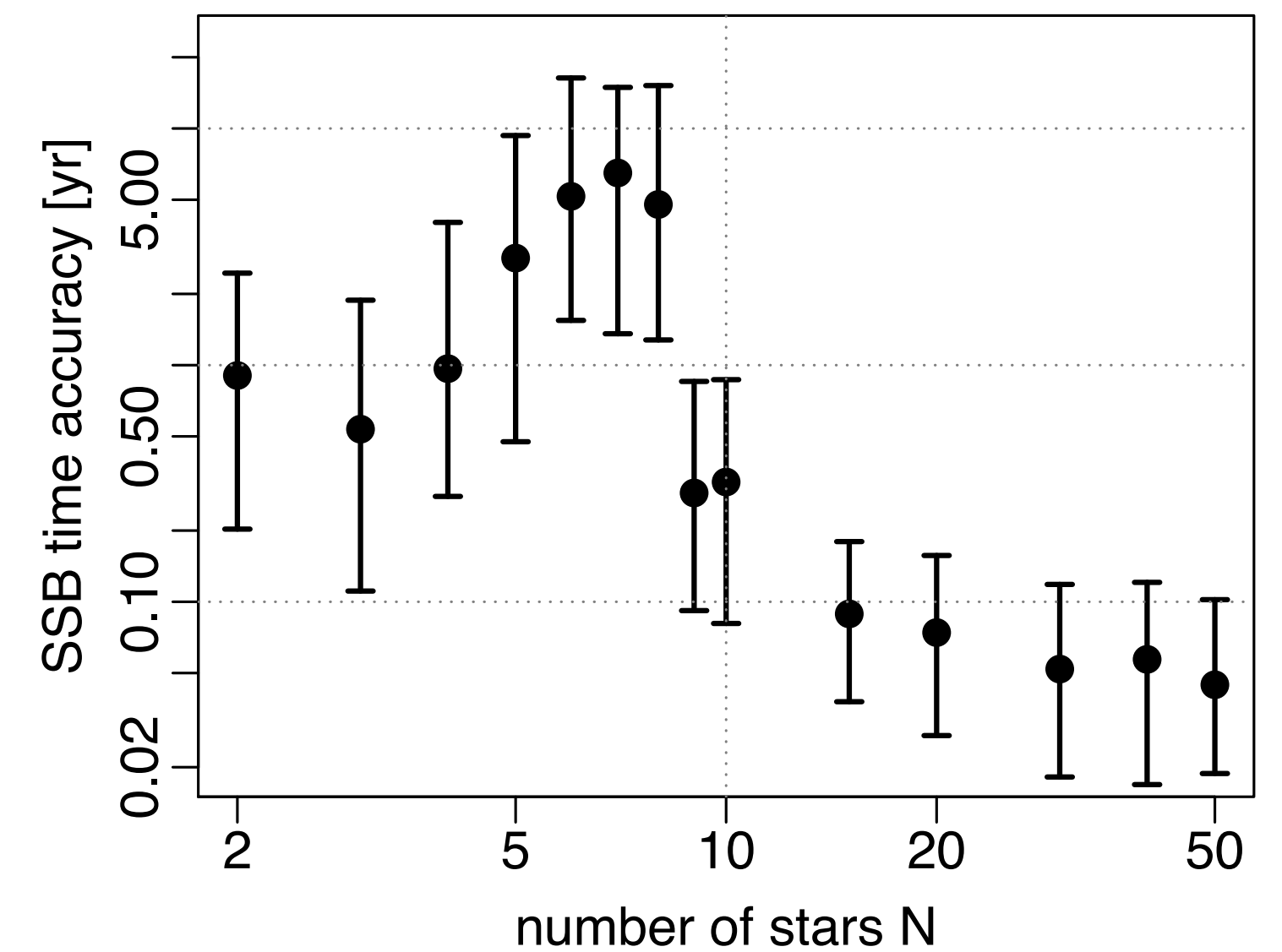
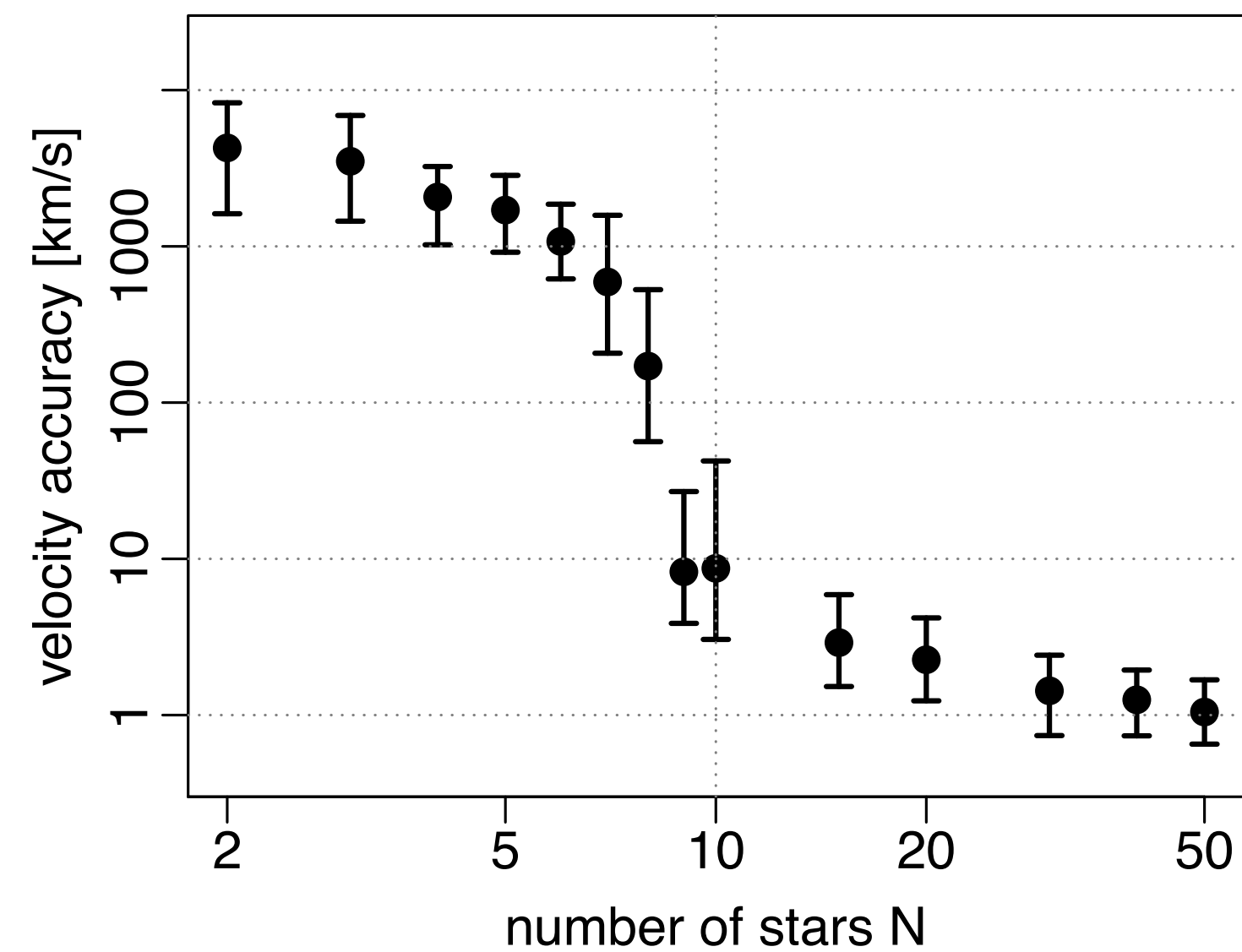
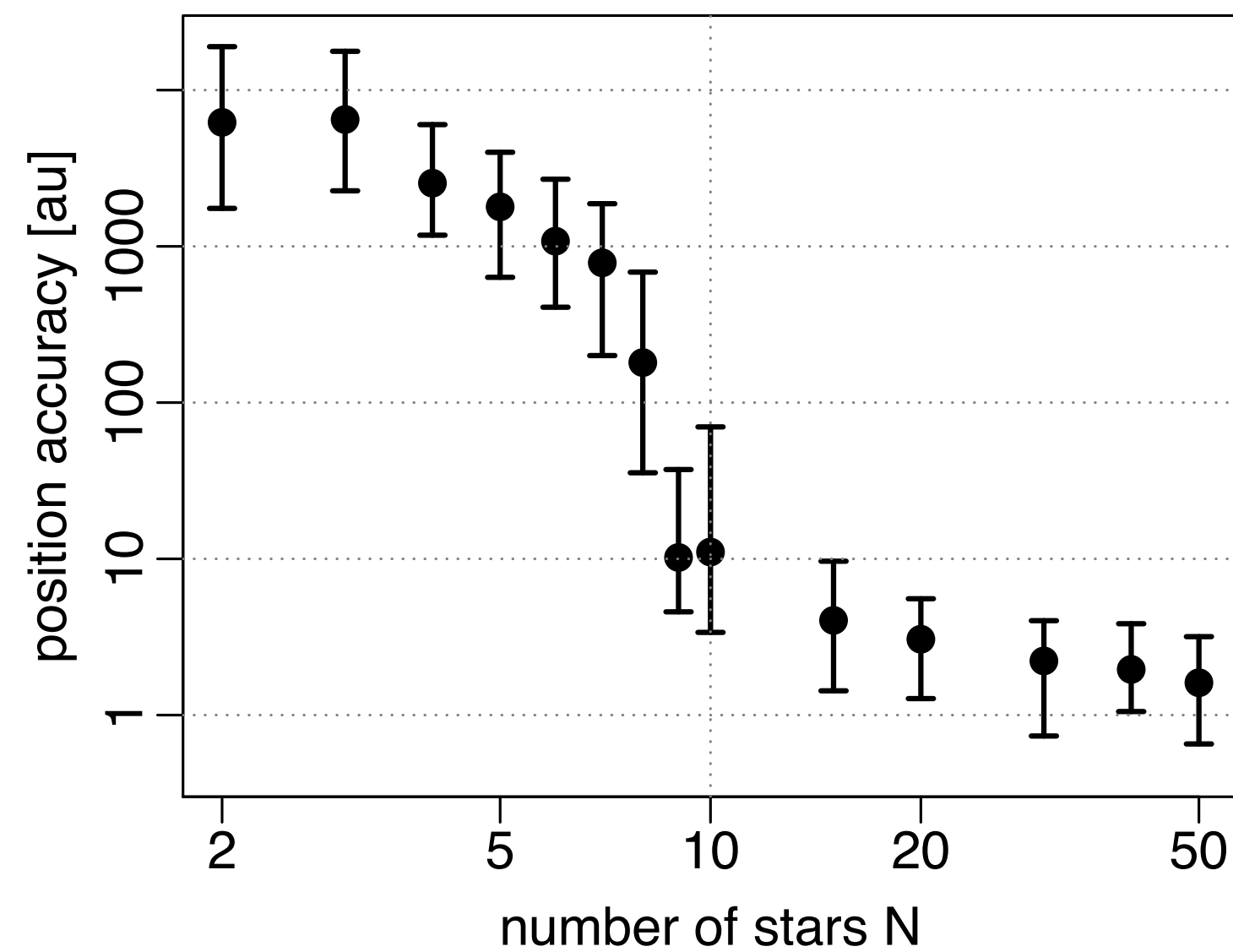
Tests on simulated data

- 20 (nearest) stars
- 1'' angular measurement accuracy
- Spacecraft placed/directed at random w.r.t. Sun
 - position: 0.1 to 10 ly away
 - velocity: 0.0 to 0.5 c
- Run iterative solution
- Evaluate for multiple setups

Inferred velocity and position accuracy



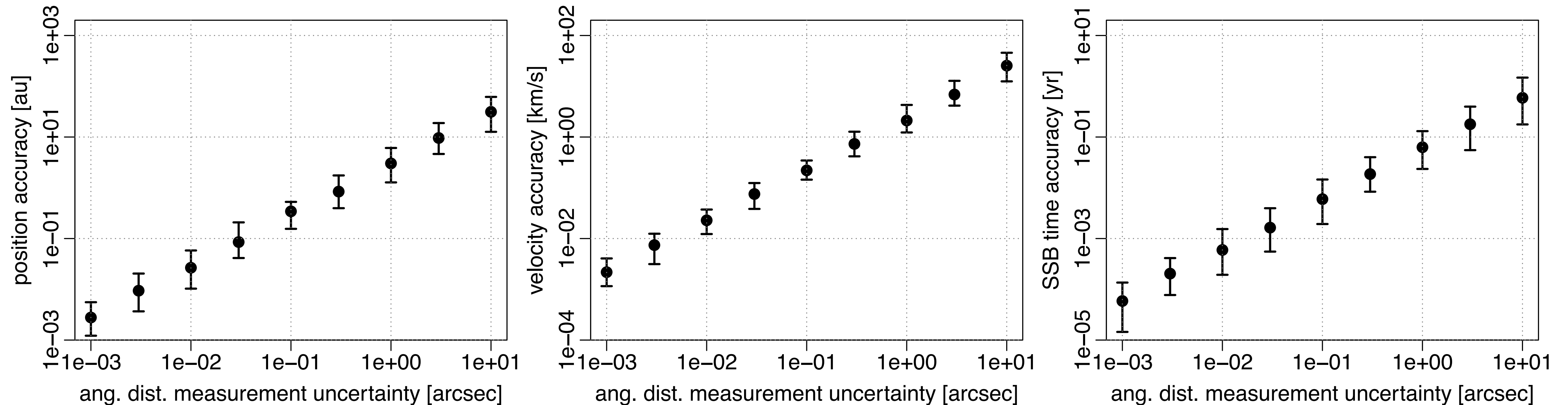
Performance: varying number of stars (1'' measurements)



Performance: varying measurement uncertainties (20 stars)



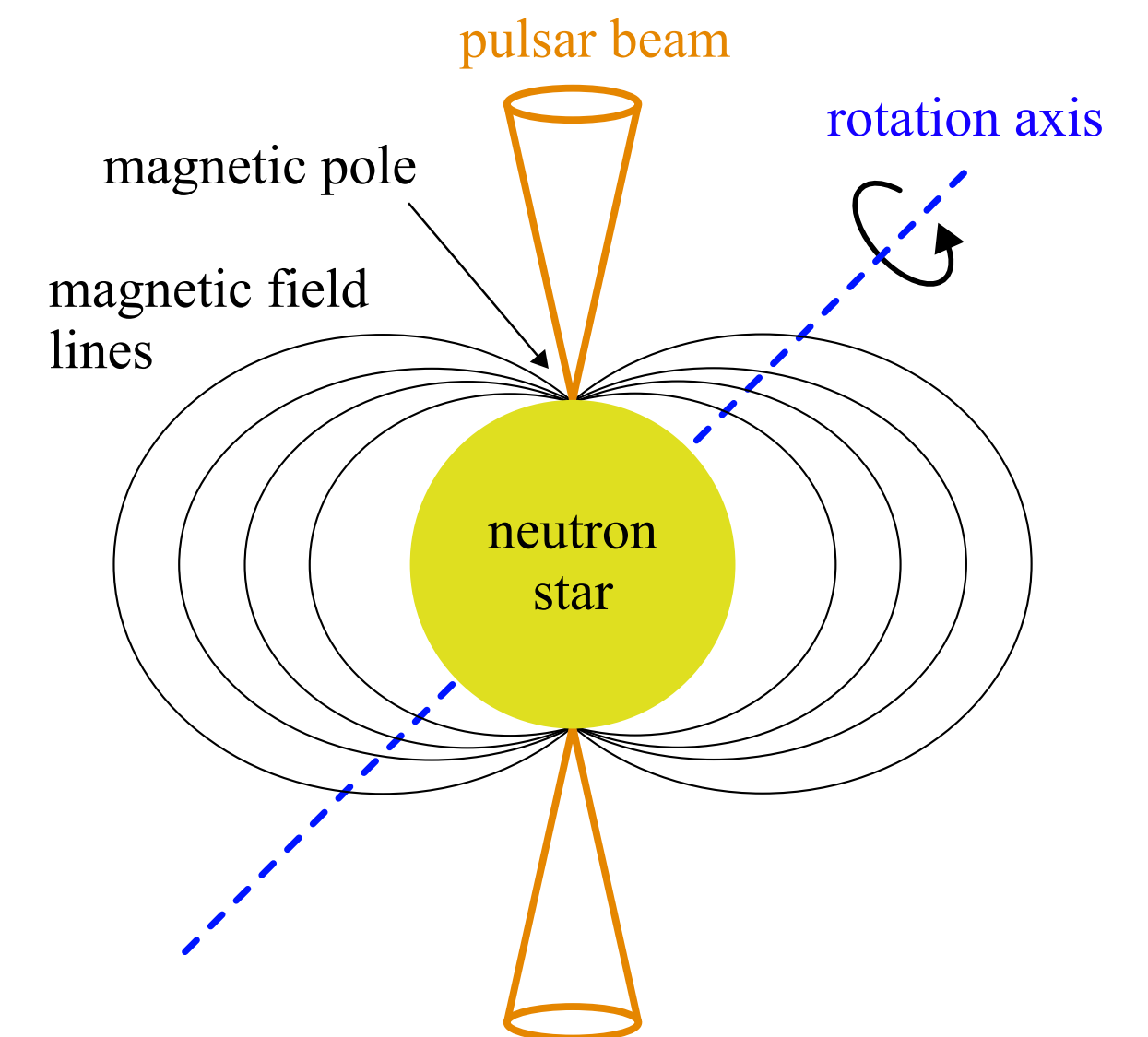
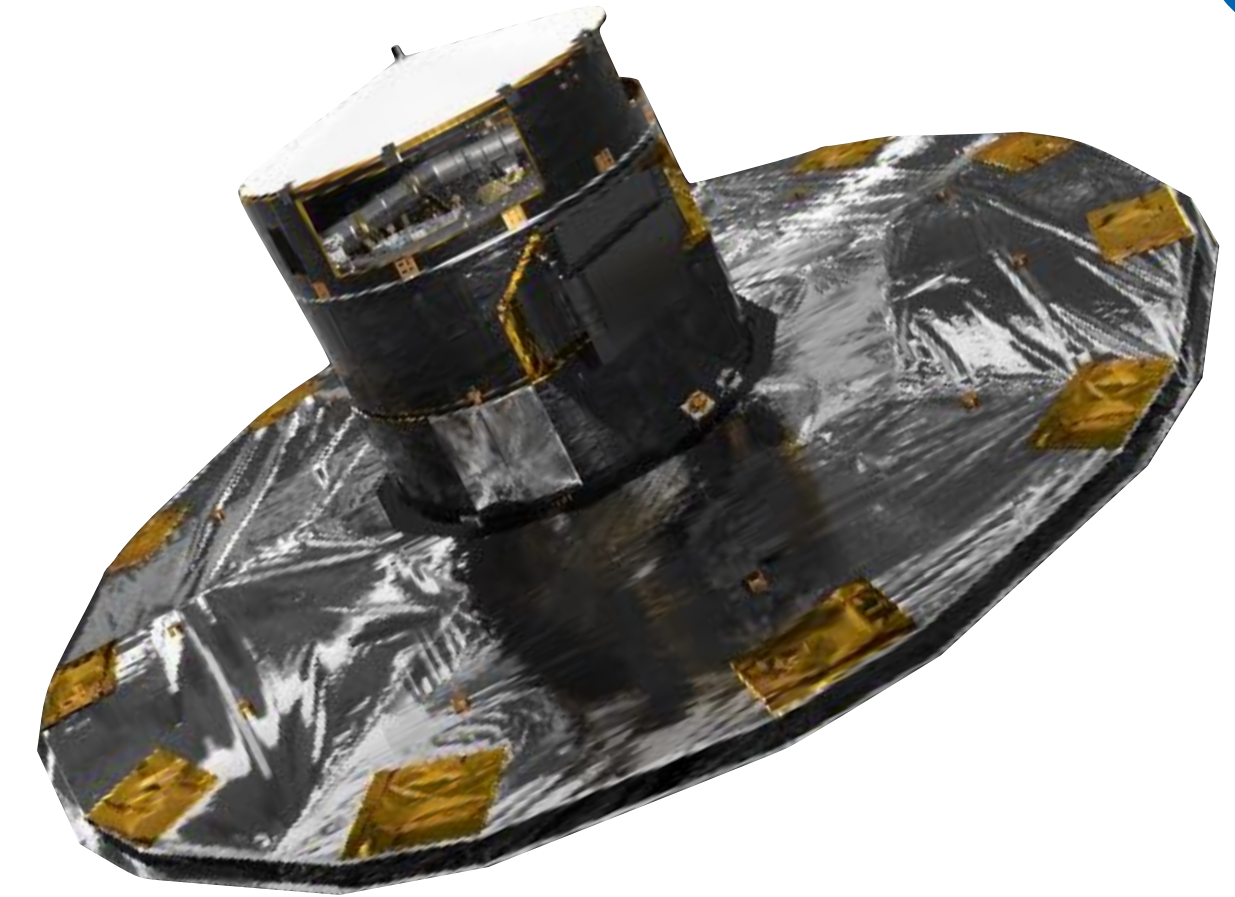
With 20 stars measured at 1 mas accuracy
can determine spacecraft position to 0.003 au, velocity to 2 m/s



Context



- Principle similar to Gaia, but in reverse
- Do not need to know SSB time
- Can use radial velocities instead (or as well)
 - ▶ need ~ 0.1 km/s Doppler measurement accuracy to determine position to 1-2 au
- Pulsar navigation gives more accurate positions, if you have a very good initialization
 - ▶ but for interstellar also need pulsar parallaxes



Conclusions



- Principle
 - ▶ Exploit parallax effect and aberration to determine spacecraft position and velocity
- Measurements required
 - ▶ Only angular separations between pairs of stars (local, instantaneous)
- Navigation accuracy
 - ▶ Using 20 stars with 1 mas accuracy, can locate spacecraft to 0.003 au, 2 m/s
- Works at large distances (100s of ly) from Sun and at relativistic velocities
- For more details see Bailer-Jones 2021 (PASP 133, 074502)